Mixed-oxide (MOX) Fuel Performance Benchmark

Summary of the Results for the Halden Reactor Project MOX Rods

Compiled by

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Work carried out within the

Expert Group on Reactor-based Plutonium Disposition (TFRPD)

Abstract

Within the framework of the Expert Group on Reactor-based Plutonium disposition (TFRPD), a fuel modelling code benchmark test for MOX fuel was initiated, with inpile irradiation data on two short MOX rods provided by the OECD Halden Reactor Project (HRP). This note summarises the provided in-pile data and fuel characteristics for the irradiation, and presents the calculation results provided by the contributors. A limited sensitivity study of the effect of the rod power uncertainty on code predictions of fuel centreline temperature and fuel pin pressure has also been performed and is included in the report.

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FOREWORD

The NEA Nuclear Science Committee has established an Expert Group that deals with the status and trends of reactor physics, nuclear fuel performance and fuel cycle issues related to the disposition of weapons-grade plutonium as mixed-oxide fuel. The objectives of the group are to provide NEA member countries with up-to-date information on, and develop consensus regarding, core and fuel cycle issues associated with weapons-grade plutonium disposition in thermal water reactors (PWR, BWR, VVER-1000 and CANDU) and fast reactors (BN-600). These issues concern core physics, fuel performance and reliability, and the capability and flexibility of thermal water reactors and fast reactors to dispose of weapons-grade plutonium in standard fuel cycles.

The activities of the NEA Expert Group on Reactor-based Plutonium Disposition are carried out in close co-operation with the NEA Working Party on Scientific Issues in Reactor Systems (WPRS), and in most cases jointly. An eminent part of these activities include benchmark studies.

This report describes the results of the first benchmark relative to MOX fuel behaviour. The corresponding experimental data have been released, compiled and reviewed for the International Fuel Performance Experiments (IFPE) Database. Other MOX fuel behaviour benchmarks being finalised or in progress are:

- Belgonucléaire and SCK.CEN PRIMO ramped MOX fuel rod performance benchmark (ongoing);
- US Department of Energy weapons-grade MOX fuel irradiation experiment irradiated at the advanced test reactor (ATR) benchmark (started);
- MOX fuel rod behaviour in fast power pulse conditions (started).

At the time of preparation of the report the following benchmarks relative to the reactor physics activities of the Expert Group were completed or in progress:

- VENUS-2 MOX core benchmarks, carried out jointly with the WPRS (completed);
- VVER-1000 LEU and MOX benchmark (completed);
- KRITZ-2 benchmarks, carried out jointly with the WPRS (completed);
- benchmark using dosimetry data from the VENUS-2, MOX core experiments (completed);
- VVER-1000 in-core self-powered neutron detector calculational benchmark (started);
- VENUS-7 weapons-grade MOX core benchmark (started).

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TABLE OF CONTENTS

Chapter 1 **INTRODUCTION**

1.1 Irradiation data

The benchmark exercise was performed on a data set provided by the OECD Halden Reactor Project (HRP) of two short MOX rods (one hollow and one solid). The rods were instrumented with fuel thermocouples (TF) and internal rod pressure transducers (PF) and irradiated in the Halden Boiling Water Reactor (HBWR). The irradiation data were provided in plain text format as local rod power (in kW/m) for each rod as a function of irradiation time comprising a total of ~626 irradiation days. Details of the fuel fabrication data as well as information about the provided irradiation data can be found in Appendix I.

The as-provided power data are shown in Figure 1 and Figure 2 for Rod 1 (solid pellets) and Rod 2 (hollow pellets) respectively. The linear heat rates for each fuel rod were given for three axial locations of the rod: bottom, middle and top. In addition the local heat rate at the axial elevation of the fuel centreline thermocouple (TF) tip were provided (denoted LHRTF). The rod average linear heat generation rate, including power uncertainty, for each rod is illustrated in Figures 3 and 4.

The measured fuel temperatures in the two rods during the irradiation are shown in Figures 5 and 6, Rod 1 (solid pellets) and Rod 2 (hollow pellets) respectively, whereas the corresponding measured rod pressures are given in Figures 7 and 8.

Of special note in the power histories is the short power spike at \sim 170 irradiation days followed by a ~50 day period with slightly higher power than before, during which notable gas release was observed in the rods as can be seen from the pressure curves in Figures 7 and 8. Also notable is the power increase towards the later stages of irradiation (at about \approx 525 days), which caused large gas release in both rods (see again Figures 7 and 8). The step-wise pressure increase seen for both rods during this period is an indication that the gap is closed at-power here.

It should be noted that the internal pressure measured in the rods during rod puncturing after irradiation was in good agreement with the end-of-life (EOL) pressure from the in-pile measurements shown in these figures.

1.2 Notes about the provided data

1.2.1 Radial power and burn-up profiles

The evolution of radial power and burn-up profiles with burn-up in the hollow and solid pellets has been calculated with the HELIOS code [1]. The results are summarised in Appendix II.

1.2.2 Considerations on the experimental uncertainty in power determination

In general, the irradiation of an instrumented fuel assembly (IFA) in the HBWR starts off with a calorimetric power calibration of the thermal power produced by the fuel in the experiment rig. The experimental uncertainty of this power calibration is generally considered to be within 5%, but somewhat dependent on the actual power produced in the rig (low power means higher uncertainty). Subsequent power calibrations at various stages of irradiation of the IFA are not always possible, and the power change with irradiation will always be reliant, to some degree, on neutronics calculations that are routinely performed for all experiments before start of irradiation.

In the case of the HRP, these neutronics calculations are performed with the HELIOS code [1]. Such calculations will always be dependent on the choice of local geometry (i.e. local surroundings of the irradiation rig) to best describe the actual physical irradiation conditions the rig is subject to throughout its time in-reactor. To evaluate the sensitivity of the HELIOS calculations with respect to the choice of surrounding fuel elements in-core, calculations have been performed for the two extreme cases actually seen by the rig during operation in-pile.

In Figures 3 and 4 the implication on average linear heat rate from the combined uncertainty of the 5% experimental uncertainty of an in-pile power calibration, and the sensitivity of neutronics calculations to changes in local surroundings are plotted for Rod 1 and Rod 2, respectively.

The estimated error increases with burn-up error from the initial 5% (power calibration uncertainty) to ca. 10% at the end of the data set.

The results of a limited sensitivity study of the effects of a $\pm 5\%$ power uncertainty on the computed fuel pin centerline temperature and fuel pin internal pressure are presented in Appendix IV.

1.3 Contributing organisations and codes

The following countries and organisations, and the respective fuel modelling codes they have used, have provided contributions to the benchmark exercise.

1.4 PIE on irradiated fuel rods

The two MOX rods were subjected to PIE after the end of irradiation. The results of the gas puncturing and mass spectrometry of the extracted gas are summarised in Appendix III.

Figure 1. Power data for Rod 1 (solid)

Figure 3. ALHR for Rod 1 (solid) including power uncertainty

Figure 5. Measured fuel centre temperature for Rod 1 (solid)

Chapter 2

BRIEF DESCRIPTION OF SOME IMPORTANT PARAMETERS OF THE CODES

2.1 Thermal conductivity of MOX

In the following subsections, a brief summary of the various codes' models for the thermal conductivity of MOX as compared to that of $UO₂$ fuel is described for each of the codes, as provided by the participants.

2.1.1 BNFL (UK) – ENIGMA

In the ENIGMA code used by BNFL, the thermal conductivity of MOX is reduced by a fixed value of 8% as compared to that of UO_2 :

$$
k_{MOX} = 0.92 \cdot k_{UO_2}
$$

where k_{MOX} is the thermal conductivity of MOX and k_{UO_2} is the thermal conductivity of UO₂.

The low temperature behaviour of the fuel thermal conductivity is modelled by an equation of the form $1/(a + bT)$, where *a* describes the phonon-impurity scattering, *b* represents the phonon-phonon scattering and *T* is the absolute temperature. The conductivity in ENIGMA is modified to an equation of the form $1/(a_0 + a_1B(1 - f) + bT)$ to represent the effects of fission products (i.e. degradation with burn-up), where *B* is the burn-up and *f* is the fraction of fission gas atoms not in solution in the matrix, i.e. the fractional sum of all gas which is either in bubbles or has been released. The same degradation is assumed for both MOX and $UO₂$ fuel [2].

2.1.2 KAERI (Korea) – COSMOS

The COSMOS code uses a thermal conductivity of MOX fuel that takes into account the amount of heterogeneity of the MOX fuel:

$$
k_{MOX} = k_{matrix} \cdot \left\{ 1 - a \cdot P_{Puk}^{\frac{2}{3}} \cdot \left[1 - 1 \cdot \left(1 + \frac{1}{a} \cdot P_{Puk}^{\frac{1}{3}} \left(\frac{k_{matrix}}{k_{Puk}} - 1 \right) \right) \right] \right\}
$$

where k_{MOX} is the thermal conductivity of MOX, k_{matrix} is the thermal conductivity of a matrix with only a small amount of Pu, k_{Puk} is the thermal conductivity of Pu rich particles, P_{Puk} is the volumetric fraction of Pu rich particles and *a* is the anisotropy factor $(a = 1$ means isotropic pore distribution).

The net effect of the above model is a reduction of 4-8% of the thermal conductivity as compared to $UO₂$ fuel [3].

2.1.3 Kurchatov Institute (RU) – FRED

The thermal and mechanical properties of LWR MOX fuel in the FRED code are based on available test data and the MATPRO-V11 material property library [4]. The table below summarises the formulation of the fuel thermal conductivity for MOX fuel used in the present calculations.

2.1.4 ORNL (USA) – FRAPCON-3 v1.3

In the version of FRAPCON-3 (v1.3, modified by ORNL for MOX applications) used by ORNL for these simulations, the MOX thermal conductivity used is a combination of the Duriez stoichiometry-dependent correlation [9] plus the burn-up degradation contained in a modified version of the NFI fuel thermal conductivity model. This combined model is described in [5].

2.1.5 ORNL (USA) – TRANSURANUS version v1m3j04

In version v1m3j04 of TRANSURANUS, used by ORNL for these simulations, a new best-estimate MOX thermal conductivity correlation has been implemented (LAMBDA=31) which is a combination of the Duriez stoichiometry-dependent correlation for fresh MOX fuel and new data (obtained via a laser flash technique at ITU) on irradiated MOX fuel with burn-up values between 20 and 45 MWd/kgHM. This model is described in [6]. Five MOX thermal conductivity correlations are available in v1m3j04, these simulations use the new best-estimate correlation (LAMBDA=31) for MOX fuel.

2.1.6 SCKCEN (Belgium) – FEMAXI-V

The FEMAXI code has three options for the thermal conductivity, which can be selected by means of the MTHCN parameter:

- MTHCN=1: Martin's model [7];
- \bullet MTHCN=2: MATPRO-11 model [4];
- \bullet MTHCN=3: Baron's model [8].

The provided results use the Baron model [MTHCN=3] for the thermal conductivity.

2.1.7 VNIINM-Bochvar (Russia) – START-3

The START-3 code used by the VNIINM for these calculations, uses a similar model as [9] (Duriez) for the thermal conductivity of (fresh) MOX fuel. In this formulation, the conductivity is constant with Pu content (a maximum Pu content of 15% is considered as the upper range for the model), but is dependent on the amount of hypostoichiometry of the fuel. The burn-up dependence of the thermal conductivity is modelled as [10] with an extra factor in the phonon term $1/(a + bT)$.

2.2 Fission gas release model

2.2.1 BNFL (UK) – ENIGMA

The fission gas release model in the ENIGMA code used by BNFL is the same as for $UO₂$, but the *fission gas generation model* is modified for MOX fuel [2].

2.2.2 KAERI (Korea) – COSMOS

The fission gas release model used for MOX fuel in the COSMOS code is the same as that used for $UO₂$, but the Pu content in the matrix is taken into account assuming that Pu-rich particles are distributed uniformly in the $UO₂$ matrix. The local fission rates inside the fuel are then calculated considering the effective Pu content in each zone [3].

2.2.3 Kurchatov Institute (RU) – FRED

For the calculation of FGR with the FRED code, the FASTGRASS model [11] is used. FASTGRASS is a mechanistic code that predicts atomic and bubble behaviour of fission gas in $UO₂$ fuel under steady-state and transient conditions. The model is employed on MOX fuel without modification. The model accounts for amount of atoms of inert gases during fission, migration of atoms, formation and migration of gas bubbles, their coalescence and solution, influence of grain boundaries and other traps on bubble behaviour. The code allows the calculation of fuel swelling vs. time as result of fission gas accumulation in the matrix and fission gas release from fuel during base-irradiation, transient and accident conditions. The area of code applicability is up to a fuel average burn-up of 50 MWd/kgU.

2.2.4 ORNL (USA) – FRAPCON-3 v1.3

The FRAPCON-3 model for FGR employed in these simulations is the Massih/Forsberg model [12]. The model is employed on MOX fuel without modification.

2.2.5 ORNL (USA) – TRANSURANUS version v1m3j04

The recommended TRANSURANUS model for FGR (FGRMOD=6), employed by ORNL for these simulations, uses the URGAS algorithm [13] with the diffusion coefficients of Hj. Matzke [14] and a constant athermal diffusion coefficient. This model option is used in conjunction with a model for intragranular fission gas release.

2.2.6 SCKCEN (Belgium) – FEMAXI-V

The FEMAXI-V code has two model options for calculating fission gas release:

- \bullet MOX=1: The fission gas is composed of 86% Xe and 14% Kr, and the FGR, which is predicted by the same model as that for $UO₂$ fuel, is multiplied by 1.3.
- MOX=2: The fission gas is composed of 87% Xe and 13% Kr, and the FGR is predicted by the same model as that for $UO₂$ fuel without modification.

Results with both models are provided.

2.2.7 VNIINM-Bochvar (Russia) – START-3

The fission gas release model utilised in the START-3 code is based on solution of a two-stage diffusion problem for fission gas in polycrystal fuel. The fission gas release is determined by diffusive flow of fission gas atoms to grain boundaries, and subsequent "quasi-diffusive" percolation of the intergranular gas to fuel-free surface. The model also takes into account the possibility of gas release by the athermal "direct recoil" and "knock-out" mechanisms [15]. Recently the model has been updated to improve its prediction at extended burn-ups [15]. Here the model is employed on MOX fuel without modification.

Chapter 3 **RESULTS**

The results from each of the contributors are presented and briefly discussed.

3.1 Fuel temperatures

3.1.1 BNFL (UK) – ENIGMA code

The results of the ENIGMA calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 9 for Rod 1 (solid) and Figure 10 for Rod 2 (hollow).

In general (i.e. for both rods) the temperatures are under-predicted when compared to the measurements from the start of irradiation. As the irradiation proceeds, the difference between predictions and measurements gradually become smaller, and in the low power period (from ~230 to ~525 irradiation days), the predicted temperatures for the hollow rod (Rod 2) reproduce the measurements very well. For the final high power period, starting at ~525 irradiation days, the predictions and measurements start to deviate again, this time with the predicted temperatures exceeding the measured values. This is particularly the case for the solid rod (Rod 1).

As noted above, the rod pressure data indicate significant fission gas release (FGR) in both rods during the high power period towards the end of the data set (from \sim 525 days) (see Section 1.1).

3.1.2 KAERI (Korea) – COSMOS

The results of the COSMOS calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 11 for Rod 1 (solid) and Figure 12 for Rod 2 (hollow).

As can be seen by the figures, two calculation results have been provided:

- 1. Inclusion of thermal conductivity recovery after \sim 525 irradiation days after which significant FGR is seen from the pressure data (see Section 1.1). These calculations are denoted "no pressure adj." in the figures (solid red curve).
- 2. Inclusion of the above thermal recovery effect above *and* an adjustment of the saturation gas atom number in the FGR model. These calculations are denoted "pressure adj." in the figures (dashed red curve).

In general both the calculations above reproduce the measured temperatures fairly well (within ca. $\pm 25^{\circ}$ C), with a slight tendency of over-prediction after ~525 days.

3.1.3 Kurchatov Institute (RU) – FRED

The results of the FRED calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 13 for Rod 1 (solid) and Figure 14 for Rod 2 (hollow).

For the solid rod (Rod 1) the FRED calculations over-estimate the measurements by $\sim 100^{\circ}$ C at the very beginning, but between $~50$ and 100 irradiation days the agreement between measurements and predictions is very good. After this the tendency for this rod is to underestimate the measurements up to the power up-rating from ~525 days, when the agreement is somewhat better.

For the hollow rod, the agreement between calculations and measurements is very good at start of life, but the calculations start to undershoot $(\sim 50^{\circ}C)$ the measurements after ~ 50 days. The discrepancy between measurements and calculations gradually decreases with time, and from \sim 300 irradiation days, the agreement between predictions and measurements is very good. The temperature increase during the power up-rating from ~525 days is also generally well reproduced in the data.

3.1.4 ORNL (USA) – FRAPCON-3 v1.3

The results of the FRAPCON-3 calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 15 for Rod 1 (solid) and Figure 16 for Rod 2 (hollow).

The agreement with measurements is very good for both the solid and the hollow rod at BOL. After ~50-100 days a tendency of under-prediction of the measurements can be seen for both rods. During the low power period from \sim 230 to \sim 525 days, however, the agreement for the hollow rod is excellent, whereas for the solid rod the temperature estimates are still below the measurements (by \sim 25-50°C). After the FGR occurs at \sim 525, the measured temperatures are over-predicted by up to \sim 125 °C for both the solid and the hollow rod.

3.1.5 ORNL (USA) – TRANSURANUS version v1m3j04

The results of the TRANSURANUS calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 17 for Rod 1 (solid) and Figure 18 for Rod 2 (hollow).

The TRANSURANUS results are very similar to the COSMOS results (see Section 3.1.2). There is a tendency of under-prediction in the first and middle part of the irradiation, while towards the end of the final high power period (after ~525 days) the temperatures are slightly over-predicted. Overall the agreement with measurements is within $\pm 25^{\circ}$ C, apart from for Rod 1 during the first part of the low power period (from \sim 230 days), when the temperatures are under-predicted by up to \sim 100 $^{\circ}$ C.

3.1.6 SCKCEN (Belgium) – FEMAXI-V

The results of the FEMAXI-V calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 19 for Rod 1 (solid) and Figure 20 for Rod 2 (hollow).

As mentioned in Section 2.1.6 above, two sets of calculations with the FEMAXI-V code (parameter MOX=1 and MOX=2) have been provided, and both results are shown in the figures.

There is little difference between the two calculation sets for both rods, and the agreement with the measurements is generally good up to the power up-rating at \sim 525 days. Some tendency of under-predicting the measurements can be seen however, particularly during the first ~230 days. After the power increase at \sim 525 days, temperatures are overestimated for both rods, but most significantly for the solid rod (Rod 1).

3.1.7 VNIINM-Bochvar (Russia) – START-3

The results of START-3 calculations of fuel temperature (calculated at the LHRTF node) are shown in Figure 21 for Rod 1 (solid) and Figure 22 for Rod 2 (hollow).

At the time of this writing, only data up to ~230 days of irradiation were provided.

For both rods the temperatures are overestimated (ca. $50-100^{\circ}$ C) at BOL but the overshoot decreases gradually and at \sim 175 days the agreement between calculations and measurements is excellent for both rods. After this a slight under-prediction is seen for the solid rod, whereas for the hollow rod the agreement also stays good here.

3.2 Fuel rod internal pressure

In this section the results of the calculations of rod internal fuel pressure is presented and compared with the measured data. The discussion focuses mainly on the above-mentioned period from ~525 days to EOL when the measured pressure rises substantially in both rods and significant FGR is seen.

3.2.1 BNFL (UK) – ENIGMA

The results of the ENIGMA calculations of rod internal pressure are shown in Figure 23 for Rod 1 (solid) and Figure 24 for Rod 2 (hollow).

For both of the rods, ENIGMA predicts a pressure step from the high power period at \sim 525 days, but the magnitude of the pressure step is lower than what is seen from the measurements. For the solid rod, the subsequent increase seen from the pressure measurements is reproduced somewhat by the calculations, but again the magnitude is lower than the measurements. For the hollow rod, the pressure steps seen in the measured data at ~570 and ~610 days (interpreted as an indication of a closed gap at power conditions) are not reproduced by the calculations. The more or less continuous pressure increase seen from the calculations would suggest that the code does not predict a closed gap for the rods at this power and burn-up.

For both rods the EOL pressure is underestimated by the calculations.

3.2.2 KAERI (Korea) – COSMOS

The results of the COSMOS calculations of rod internal pressure are shown in Figure 25 for Rod 1 (solid) and Figure 26 for Rod 2 (hollow).

For the COSMOS calculations, the EOL pressure for the hollow rod (Rod 2) is predicted excellently with Option 2 (see Section 3.1.2 above) for the fission gas release calculations, whereas Option 1 underestimates the EOL pressure. For the solid rod, both options predict an EOL pressure below what was measured. As for the ENIGMA code discussed above, the step-wise pressure increase during the final high power period is not reproduced by the calculations, and a more or less continuous

pressure increase is predicted. However, during the period from ~525 to ~570 days, when the measured pressure in the solid rod also shows a steady increase, the calculated rates of pressure increase (for both options) are close to parallel to the increase rate of the measured pressure (Figure 25).

3.2.3 Kurchatov Institute (RU) – FRED

The results of the FRED calculations of rod internal pressure are shown in Figure 27 for Rod 1 (solid) and Figure 28 for Rod 2 (hollow).

For the high-power period after ~525 days, the FRED calculations predict pressure increase and FGR, but smaller than what was measured in-pile. The step-wise pressure increase in the measured pressures during this period is not particularly visible in the calculations. For both rods, the EOL pressure is underestimated in the calculations, mostly so for the solid rod (Rod 1).

3.2.4 ORNL (USA) – FRAPCON-3 v1.3

The results of the FRAPCON-3 calculations of rod internal pressure are shown in Figure 29 for Rod 1 (solid) and Figure 30 for Rod 2 (hollow).

The code predicts a pressure step at ~525 days for both the hollow and the solid rod, but for both rods the step is smaller than the measured increase. For both rods, a more step-wise pressure increase after the high power period starts at \sim 525 is calculated by the code than the results from both ENIGMA and COSMOS showed.

The pressure at EOL is best met for the hollow rod, but for both rods the predicted EOL pressure is somewhat below the measurements.

3.2.5 ORNL (USA) – TRANSURANUS version v1m3j04

The results of the TRANSURANUS calculations of rod internal pressure are shown in Figure 31 for Rod 1 (solid) and Figure 32 for Rod 2 (hollow).

For the high-power period after ~525 days, TRANSURANUS results show a small pressure increase, but this is much lower than what the in-pile measurements show. The step-wise pressure increase in the measured pressures during this period is not seen in the calculations. For both rods, the EOL pressure is underestimated in the calculations, mostly so for the solid rod (Rod 1).

3.2.6 SCKCEN (Belgium) – FEMAXI-V

The results of the FEMAXI-V calculations of rod internal pressure are shown in Figure 33 for Rod 1 (solid) and Figure 34 for Rod 2 (hollow).

As for the temperature calculations (see Section 3.1.5), two sets of calculations are provided, where the difference is in the FGR model. For the hollow rod, the difference between the two models is insignificant, and the EOL pressure is underestimated quite a bit in both cases. For the solid rod (Rod 1) however, the two models yield different results. In the case of this rod, the highest pressure increase is predicted by the MOX=1 option, coming relatively close to the measurements in predicting the EOL pressure.

As for the COSMOS and ENIGMA code, the pressure development after ~525 days shows a more or less steady increase for the solid rod (both for the MOX=1 and MOX=2 option), and does not reproduce very well the stair-steps seen in the measured pressure for this rod.

3.2.7 VNIINM-Bochvar (Russia) – START-3

The results of the START-3 calculations of rod internal pressure are shown in Figure 35 for Rod 1 (solid) and Figure 36 for Rod 2 (hollow).

As for the temperature data only data up to ~225 days are provided at the time of this writing. During this irradiation period, there is not much development in the measured pressure, and it is thus difficult to make any judgement of how well the code is able to reproduce the in-pile behaviour of these rods.

Figure 9. Fuel temperature for Rod 1 (solid) calculated by ENIGMA (BNFL)

TFRPD results Rod No. 1 (solid): Fuel temperature (°C)

Figure 10. Fuel temperature for Rod 2 (hollow) calculated by ENIGMA (BNFL)

TFRPD results Rod No. 2 (hollow): Fuel temperature (°C)

Figure 11. Fuel temperature for Rod 1 (solid) calculated by COSMOS (KAERI)

TFRPD results Rod No. 2 (hollow): Fuel temperature (°C)

Figure 13. Fuel temperature for Rod 1 (solid) calculated by FRED (Kurchatov)

Figure 14. Fuel temperature for Rod 2 (hollow) calculated by FRED (Kurchatov)

TFRPD results Rod No. 2 (hollow): Fuel temperature (C)

TFRPD results Rod No. 1 (solid): Fuel temperature (C)

TFRPD results Rod No. 2 (hollow): Fuel temperature (C)

Figure 17. Fuel temperature for Rod 1 (solid) calculated by TRANSURANUS (ORNL)

Figure 18. Fuel temperature for Rod 2 (hollow) calculated by TRANSURANUS (ORNL)

TFRPD results Rod No. 2 (hollow): Fuel temperature (°C)

Figure 19. Fuel temperature for Rod 1 (solid) calculated by FEMAXI-V (SCK•CEN)

Figure 20. Fuel temperature for Rod 2 (hollow) calculated by FEMAXI-V (SCK \bullet CEN)

TFRPD results Rod No. 2 (hollow): Fuel temperature (C)

Figure 21. Fuel temperature for Rod 1 (solid) calculated by START-3 (VNIINM)

TFRPD results Rod No. 2 (hollow): Fuel temperature (°C)

Figure 23. Internal rod pressure for Rod 1 (solid) calculated by ENIGMA (BNFL)

TFRPD results Rod No. 1 (solid): Rod pressure (bar)

TFRPD results Rod No. 2 (hollow): Rod pressure (bar)

Figure 25. Internal rod pressure for Rod 1 (solid) calculated by COSMOS (KAERI)

Figure 26. Internal rod pressure for Rod 2 (hollow) calculated by COSMOS (KAERI)

Figure 27. Internal rod pressure for Rod 1 (solid) calculated by FRED (Kurchatov)

TFRPD results Rod No. 1 (solid): Rod pressure (bar)

TFRPD results Rod No. 2 (hollow): Rod pressure (bar)

Figure 29. Internal rod pressure for Rod 1 (solid) calculated by FRAPCON-3 (ORNL)

TFRPD results Rod No. 1 (solid): Rod pressure (bar)

TFRPD results Rod No. 1 (solid): Rod pressure (bar)

Figure 32. Internal rod pressure for Rod 2 (hollow) calculated by TRANSURANUS (ORNL)

TFRPD results Rod No. 2 (hollow): Rod pressure (bar)

Figure 33. Internal rod pressure for Rod 1 (solid) calculated by FEMAXI-V (SCK•CEN)

Figure 34. Internal rod pressure for Rod 2 (hollow) calculated by FEMAXI-V (SCK.CEN)

TFRPD results Rod No. 2 (hollow): Rod pressure (bar)

TFRPD results Rod No. 1 (solid): Rod pressure (bar)

Figure 36. Internal rod pressure for Rod 2 (hollow) calculated by START-3 (VNIINM)

TFRPD results Rod No. 2 (hollow): Rod pressure (bar)

REFERENCES

- [1] Skardhamar, T., U. Kasemeyer, "Introduction of HELIOS to the Halden Reactor Project", *Studsvik Scandpower European CMS/FMS Users Group Meeting*, Madrid (1998).
- [2] Palmer, I., G. Rossiter, R. White, *Development and Validation of the ENIGMA Code for MOX Fuel Performance Modelling*, IAEA-SM-358/20.
- [3] Lee, B-H., Y-H. Koo, J-S. Cheon, J-Y. Oh, D-S. Sohn, "COSMOS Code Benchmark by Using a Solid and Hollow MOX Fuel Behaviour", Contribution to the MOX Benchmark Test (2003).
- [4] Hagrman, D.L., G.A. Reymann, R.E. Mason, *MATPRO-Version 11 (Revision 1): A Handbook of Material Properties for Use in the Analysis of Light Water Reactor Fuel Rod Behavior*, NUREG/CR-0497, Rev 1, February 1980.
- [5] Lanning, D.D., C.E. Beyer, "Revised Fuel Thermal Conductivity for NRC Fuel Performance Codes", Transactions of the American Nuclear Society, 2002 Annual Meeting, Hollywood, Florida, Vol. 86, pp. 285-287, June 2002.
- [6] Schubert, A., P.V. Uffelen, J.V.D. Laar, M. Sheindlin, L. Ott, "Present Status of the MOX Version of the TRANSURANUS Code (F2.5)", *Enlarged Halden Programme Group Meeting on High Burn-up Fuel Performance, Safety and Reliability*, Rica Park Hotel, Sandefjord, Norway, 9-14 May 2004.
- [7] Martin, D.G., "A Re-appraisal of the Thermal Conductivity of $UO₂$ and Mixed (U,Pu) Oxide Fuels", *J. Nucl. Mat*., 110, pp. 73-94 (1982).
- [8] Baron, D., "About the Modelling of Fuel Thermal Conductivity Degradation at High-burn-up Accounting for Recovering Process with Temperature", *Proc Sem. Thermal Performance of High-burn-up LWR Fuel*, Cadarache, France, 3-6 March 1998, pp. 129-143.
- [9] Duriez, C., J-P. Alessandri, T. Gervais, Y. Philipponneau, "Thermal Conductivity of Hypostoichiometric Low Pu Content (U,Pu)O_{2-x} Mixed Oxide", *J. Nucl. Mat.*, 277, pp. 143-158 (2000).
- [10] Wiesenack, W., T. Tverberg, "Thermal Performance of High Burn-up Fuel", *ANS Topical Meeting on LWR Fuel Performance* (2000).
- [11] Rest, J., S.A. Zawadzki, *FASTGRASS: A Mechanistic Model for the Prediction of Xe, I, Cs, Te, Ba, and Sr Release from Nuclear Fuel Under Normal and Severe Accident Condition*, NUREG/CR-5840 ANL-92/3, September 1992.
- [12] Lanning, D.D., C.E. Beyer, C.L. Painter, *FRAPCON-3: Modifications to Fuel Rod Material Properties and Performance Models for High-burn-up Application*, NUREG/CR-6534, Chapter 2, October 1997.
- [13] Lassmann, K., H. Benk, "Numerical Algorithms for Intragranular Fission Gas Release", *J. Nucl. Mater*., 280, pp. 127-135 (2000).
- [14] Matzke, Hj., "Gas Release Mechanisms in UO₂ A Critical Overview", *Radiation Effects*, 53, pp. 219-242 (1980).
- [15] Bibilashvili, Yu.K., A.V. Medvedev, G.A. Khvostov, S.M. Bogatyr, L.V. Korystin, "Development of the Fission Gas Behaviour Model in the START-3 Code and its Experimental Support", *Seminar on Fission Gas Behaviour in Water Reactor Fuels*, Cadarache, France, 26-29 September 2000.

Appendix I

SPECIFICATION FOR A SOLID AND HOLLOW MOX PELLET BEHAVIOUR BENCHMARK

Revision 4 – 2 June 2003

(Revision 3 – 19 November 2002)

(Revision $2 - 9$ October 2001 – ~235 irradiation days)

(Revision 1 – 6 June 2000)

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This version includes data up to ~626.5 irradiation days and the total free volume for the 2 rods.

Local heat rates:

ROD2.dat

Appendix II

RADIAL POWER AND BURN-UP PROFILES

Figure A.1. Relative power and burn-up distribution in Rod 1 (solid pellet) as function of radius

Figure A.2. Relative power and burn-up distribution in Rod 2 (hollow pellet) as function of radius

Figure A3. Relative power and burn-up distribution in Rod 1 (solid pellet) as function of burn-up

Relative Power and Burnup Distribution in solid pellet

Table A.1. Radial power distribution for Rod 1 (solid pellet)

Table A.2. Radial power distribution for Rod 2 (hollow pellet)

	Radially averaged burn-up [MWd/kgOX]																
	0.00	0.42	0.85	1.69	3.39	5.08	6.77	8.46	11.84	15.21	16.11	21.19	23.05	23.88	24.72	28.05	31.38
Radius																	
[mm]	Relative power																
1.5550	0.8172	0.8196	0.8215	0.8264	0.8364	0.8466	0.8568	0.8669	0.8870	0.9063	0.9128	0.9369	0.9438	0.9470	0.9506	0.9625	0.9710
2.0070	0.8363	0.8385	0.8403	0.8449	0.8541	0.8634	0.8727			0.8819 0.9000 0.9172	0.9231	0.9438	0.9487	0.9514	0.9544	0.9643	0.9710
2.3740	0.8681	0.8700	0.8715	0.8754	0.8831	0.8909	0.8985	0.9060	0.9205	0.9339	0.9382	0.9533	0.9567	0.9586	0.9607	0.9672	0.9711
2.6910	0.9054	0.9069	0.9080		0.9110 0.9168	0.9225	0.9281		0.9335 0.9436	0.9525	0.9551	0.9638	0.9653	0.9662	0.9673	0.9702	0.9711
2.9750	0.9476	0.9484	0.9492	0.9510	0.9545	0.9578	0.9609	0.9638	0.9687	0.9725	0.9729	0.9748	0.9738	0.9738	0.9739	0.9731	0.9712
3.2340	0.9948	0.9950	0.9951	0.9955	0.9961	0.9965	0.9966	0.9965	0.9956	0.9937	0.9924	0.9869	0.9844	0.9834	0.9823	0.9779	0.9733
3.4740	1.0487	1.0481	1.0475	1.0462	1.0433	1.0401	1.0368		1.0332 1.0256	1.0175	1.0140	1.0012	0.9966	0.9947	0.9926	0.9851	0.9783
3.6980	1.1108	1.1091	1.1078		1.1043 1.0972	1.0899	1.0826	1.0752	1.0603	1.0457	1.0403	1.0209	1.0155	1.0128	1.0100	1.0001	0.9924
3.9090	1.1851	1.1823	1.1799	1.1740	1.1621	1.1503	1.1386	1.1272	1.1054	1.0853	1.0795	1.0566	1.0511	1.0482	1.0451	1.0355	1.0296
4.1100	1.2861	1.2822	1.2791		1.2714 1.2564 1.2419 1.2283						1.2157 1.1933 1.1753 1.1717 1.1618 1.1642			1.1639	1.1631	1.1641	1.1710

Table A.3. Radial burn-up distribution for Rod 1 (solid pellet)

Table A.4. Radial burn-up distribution for Rod 2 (hollow pellet)

Appendix III

RESULTS OF GAS PUNCTURING AND GAS ANALYSIS

Table A.5. Results of gas puncturing

* Corrected 17 June 2005 (error in previous version of this report; old value: 3.20).

Table A.6. Results of mass spectrometry of extracted gas

Rod		1 (solid)	2 (hollow)		
83Kr	[mg]	3.81	3.67		
${}^{84}\text{Kr}$	[mg]	7.41	7.17		
85 Kr	[mg]	1.50	1.44		
${}^{86}\text{Kr}$	[mg]	11.63	11.23		
Total Kr	[mg]	24.35	23.51		
$\overline{^{128}}$ Xe	[mg]	0.30	0.30		
130 Xe	[mg]	0.86	0.86		
131 Xe	[mg]	67.28	64.57		
132 Xe	[mg]	128.10	124.20		
133 Xe	[mg]	1.23	1.16		
134 Xe	[mg]	169.40	163.70		
$^{136}\mathrm{Xe}$	[mg]	253.80	246.70		
Total Xe	[mg]	620.97	601.49		

Table A.7. ORIGEN calculations of inventory of fission gases at end of life

Table A.8. Fission gas release from fuel

Rod				
Total Kr released [mg]	5.66	4.55		
Total Xe released	106.14	85.96		
Total $Kr + Xe$ released	111.80	90.51		
Total Kr generated	24.35	23.51		
Total Xe generated	620.97	601.49		
Total $Kr + Xe$ generated	645.32	625.00		
Fraction $Kr + Xe$ released	0.1732	0.1448		

Appendix IV

RESULTS OF "LIMITED" SENSITIVITY STUDY FOR THE SOLID AND HOLLOW MOX PELLETS BEHAVIOUR BENCHMARK

In order to get a picture of the effect of the rod power uncertainty on the calculated fuel centreline temperature and fuel pin pressure, a sensitivity study was performed at the ORNL. The codes used in the study were FRAPCON-3 and TRANSURANUS.

In this study both code-to-code and code-to-experiment comparisons are made. In the following, the results of the sensitivity study are presented.

The basis for the sensitivity study are the uncertainties in the linear heat generation rates which show an uncertainty varying from 5% initially to 10% at EOL. In the ORNL calculations, however, a 5% uncertainty is assumed throughout the irradiation.

The main features and outcomes are listed below.

Objectives of the sensitivity study

- \bullet Determination the effect of the rod power uncertainty on the calculated fuel centreline temperature and fuel pin pressure.
- \bullet A code-to-code comparison of FRAPCON-3 and TRANSURANUS as well as code-toexperiment comparisons for each code.
- Study of the effect of fuel thermal conductivity correlations.

General code-to-code-to-experiment conclusions

- - The TRANSURANUS temperature predictions compare more favourably with data than FRAPCON-3 (Figures A5, A6, A15 and A16).
- The FRAPCON-3 pressure predictions are much better than TRANSURANUS (Figures A7, A8, A17 and A18).
- TRANSURANUS does not predict significant fission gas release (throughout the irradiation); therefore, the code significantly underpredicts the rod pressure response.
- \bullet The FRAPCON-3 MOX thermal conductivity correlation [5] and the ITU correlation [6] within the same framework (i.e. code) yield essentially equivalent fuel centreline temperatures (figure not shown).
- \bullet Code-to-code modelling differences (for example, fuel densification and swelling, fuel relocation, FGR models) account for the significant differences in the code-to-code predictions.

General conclusions on power uncertainty effects

- \bullet Without the FGR feedback, an uncertainty of $\pm 5\%$ in the power generation results in an uncertainty of \leq 5% in the calculated fuel centreline temperature (see FRAPCON-3 results, Figures A11 and A12 at times <175 days, or TRANSURANUS results, Figures A19 and A20 throughout the irradiation).
- \bullet Realistically, the temperature differences (produced by $\pm 5\%$ power uncertainty) significantly affect the fission gas release models and FGR which in turn affects the pellet-to-clad gap conductance. This can result in:
	- -10 to +15% range on the computed fuel centreline temperature (see Figures A11 and A12);
	- -30 to +40% range on the computed fuel rod pressure (see Figures A13 and A14).
- \bullet The inclusion of power uncertainties does produce calculated temperatures and pressures (FRAPCON-3) which envelope the experimental data (see Figures A5 through A8).

Figure A.5. FRAPCON-3 Rod 1 (solid) centreline temperature

Figure A.6. FRAPCON-3 Rod 2 (hollow) centreline temperature

Figure A.7. FRAPCON-3 Rod 1 (solid) internal pressure

Figure A.8. FRAPCON-3 Rod 2 (hollow) internal pressure

Figure A.9. FRAPCON-3 Rod 1 (solid) calculated fission gas release

Figure A.10. FRAPCON-3 Rod 2 (hollow) calculated fission gas release

Figure A.11. Effect of 5% power variation on Rod 1 (solid) FRAPCON-3 calculated temperature

Figure A.12. Effect of 5% power variation on Rod 2 (hollow) FRAPCON-3 calculated temperature

Figure A.13. Effect of \pm 5% power variation on Rod 1 (solid) FRAPCON-3 calculated pressure

Figure A.14. Effect of 5% power variation on Rod 2 (hollow) FRAPCON-3 calculated pressure

Figure A.15. TRANSURANUS Rod 1 (solid) centreline temperature

Figure A.16. TRANSURANUS Rod 2 (hollow) centreline temperature

Figure A.17. TRANSURANUS Rod 1 (solid) internal pressure

Figure A.18. TRANSURANUS Rod 2 (hollow) internal pressure

Figure A.19. Effect of 5% power variation on Rod 1 (solid) TRANSURANUS calculated temperature

Figure A.20. Effect of 5% power variation on Rod 2 (hollow) TRANSURANUS calculated temperature

